

Spaceborne Polarimetric SAR Interferometry: Performance Analysis and Mission Concepts

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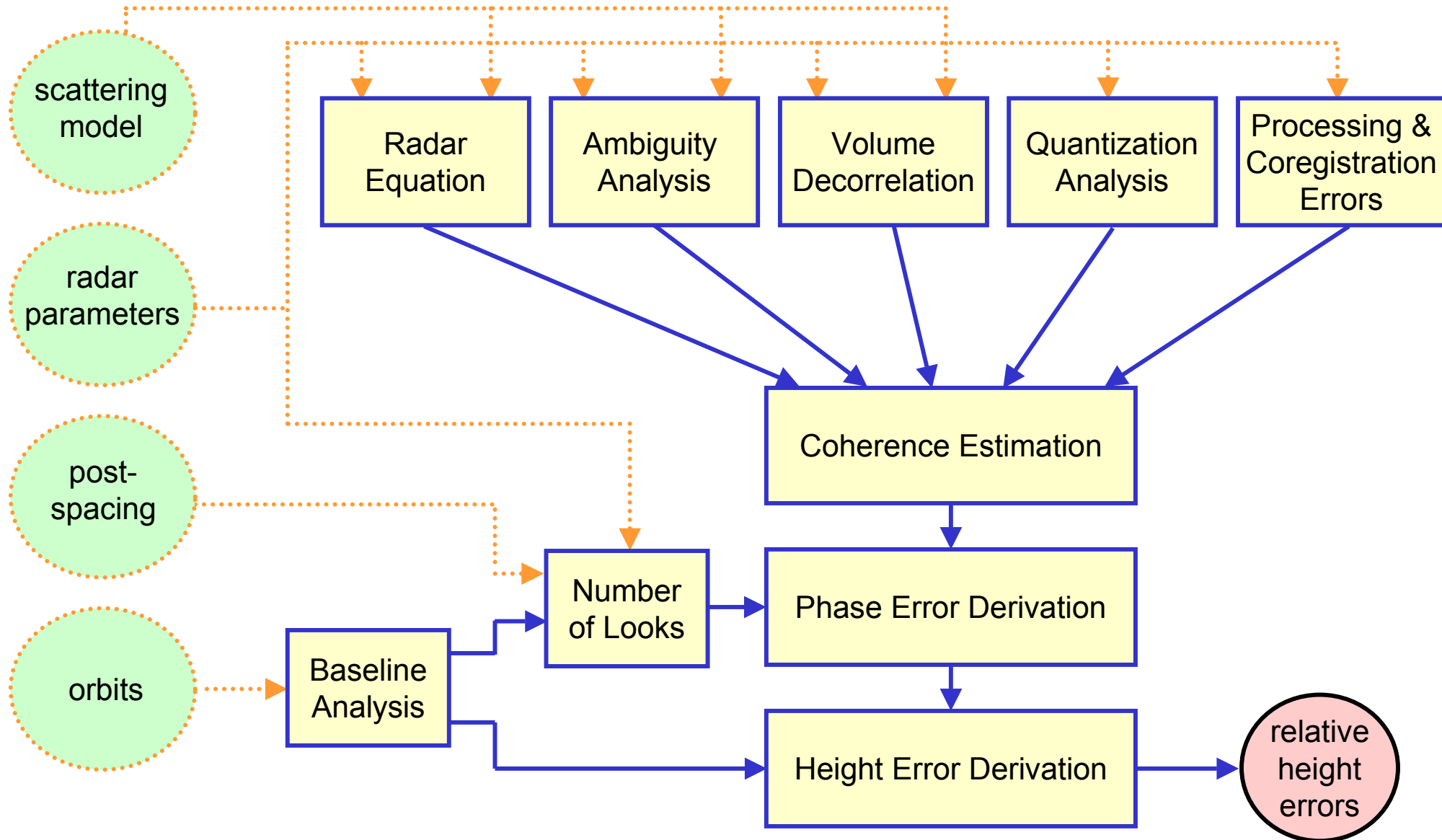
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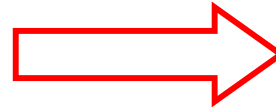
-
- *Outline of Performance Analysis*
 - *Repeat Pass Mission Scenarios*
 - *Single Pass Mission Scenarios*
 - *Critical Issues*
-



Data for Scots Pine Forest (from S. Cloude et al., 2004)

Parameter	Value
Height	18 m
Incident Angle	45°
Stem Density	0.055 stems/m ²
Extinction	0.3 dB/m
Ground to Volume Ratios	$\mu_{hh} = -12$ dB $\mu_{hv} = -31$ dB $\mu_{vv} = -18$ dB $\mu_{max} = -7$ dB $\mu_{min} = -31$ dB
Scattering Coefficients	$\sigma_{hh} = -4$ dB m ² /m ² $\sigma_{hv} = -11$ dB m ² /m ² $\sigma_{vv} = -8$ dB m ² /m ²

*Adaptation
to Satellite
Case*




Reference Scenario for Performance Evaluation

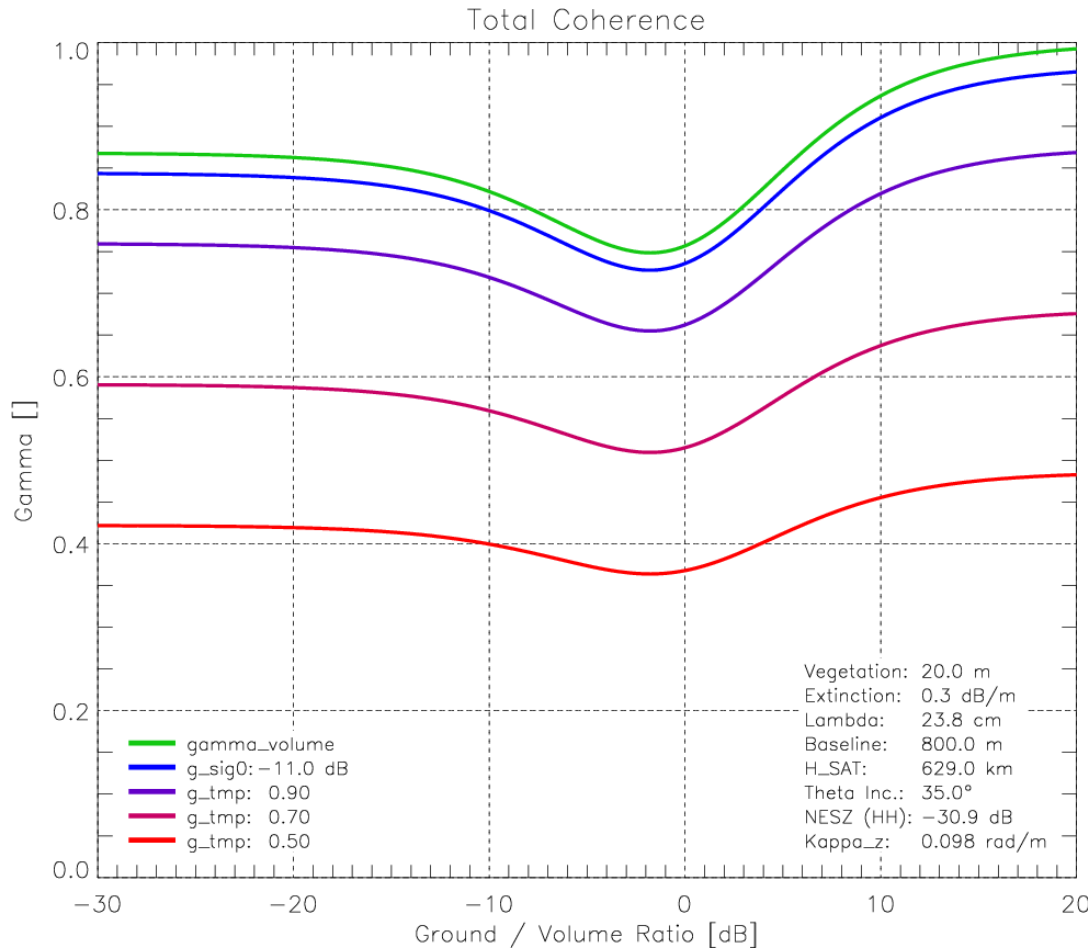
Parameter	Value
Incident Angle	35°
Volume Height	20 m
Sigma Nought	> -11 dBm ² /m ²
Minimum Ground to Volume Ratio	-26 dB
Maximum Ground to Volume Ratio	-2 dB
Extinction	0.3 dB/m

Assumptions:

- Contributions from ground will become stronger for lower incident angles
- Equal scattering coefficients of -11 dB m²/m² for all polarisations (worst case assumption)

Parameter	Value	
Wavelength	0.238 m	
Orbit Height	629 km	
Chirp Bandwidth	80 MHz	
Peak Transmit Power	6400 W (160x40W)	
Duty Cycle	3,5 % (7 % / 2)	
Tx Losses	1.3 dB	
Noise Figure	2.5 dB	
Proc. Losses	1 dB	
Losses Across Swath	< 3 dB	
Atmospheric Losses	1 dB	
Antenna Size (Tx , Rx)	11 m x 2.86 m	
Co-Registration Accuracy	1/10 pixel	
Quantisation	4 bit (BAQ)	
Access Range (quad pol mode)	20° - 36°	

Total Coherence: $\gamma_{tot} = \gamma_{SNR} \cdot \gamma_{vol} \cdot \gamma_{tmp} \cdot \gamma_{proc} \cdot \gamma_{quant} \cdot \gamma_{amb} \cdot \gamma_{geo} \cdot \gamma_{az}$



- γ_{vol} only
- γ_{tot} for $\gamma_{tmp} = 1.0$
- γ_{tot} for $\gamma_{tmp} = 0.9$
- γ_{tot} for $\gamma_{tmp} = 0.7$
- γ_{tot} for $\gamma_{tmp} = 0.5$

Coherence →

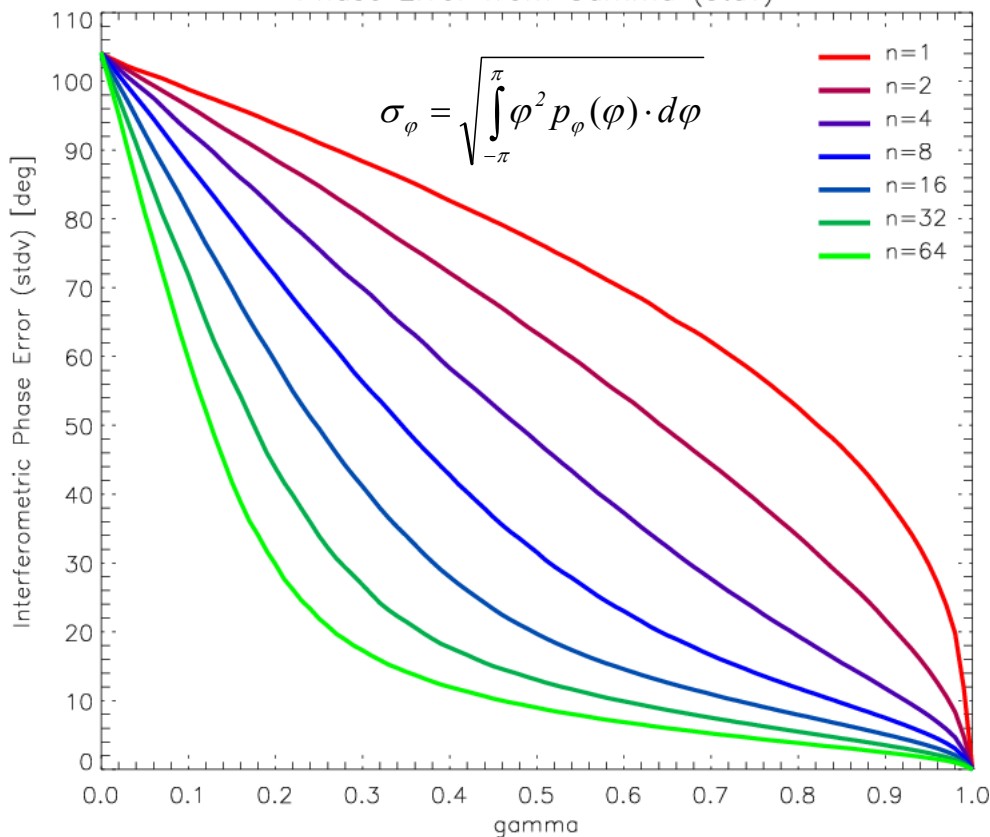
**Number
of Looks** →

post -
spacing

system & orbit
parameters

Derivation of Phase Errors

Phase Error from Gamma (stdv)



$$\sigma_{\varphi} = \sqrt{\int_{-\pi}^{\pi} \varphi^2 p_{\varphi}(\varphi) \cdot d\varphi}$$

$$p_{\varphi}(\varphi) = \frac{\Gamma(n + 1/2)(1 - \gamma^2)^n \gamma \cos \varphi}{2\sqrt{\pi} \Gamma(n)(1 - \gamma^2 \cos^2 \varphi)^{n+1/2}} + \frac{(1 - n^2)^n}{2\pi} F(n, 1; 1/2; \gamma^2 \cos^2 \varphi)$$

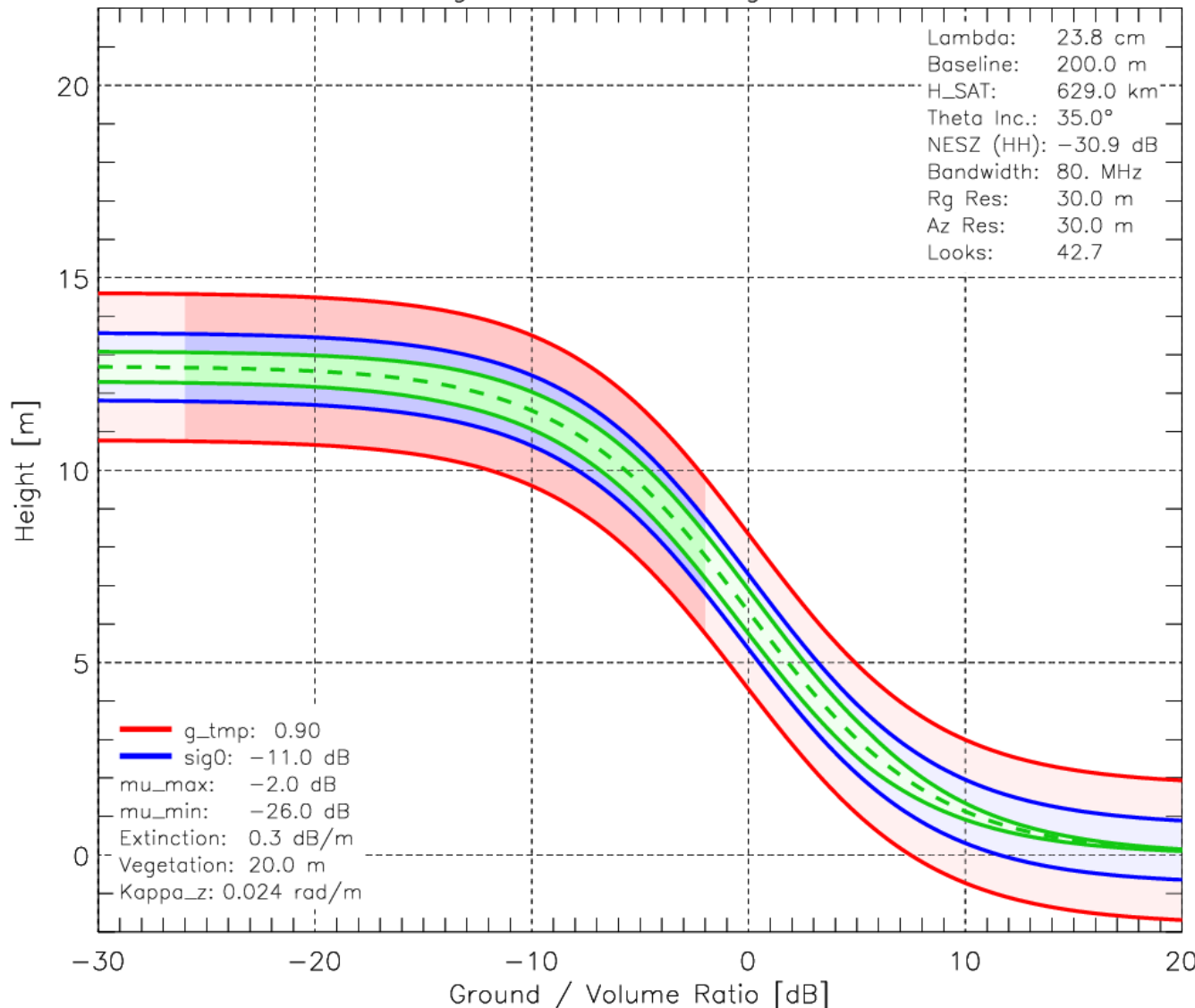
Phase to Height Conversion

$$\Delta h = \frac{\lambda r \sin(\theta_i)}{4\pi B_{\perp}} \cdot \Delta \varphi$$

interferometric
baseline

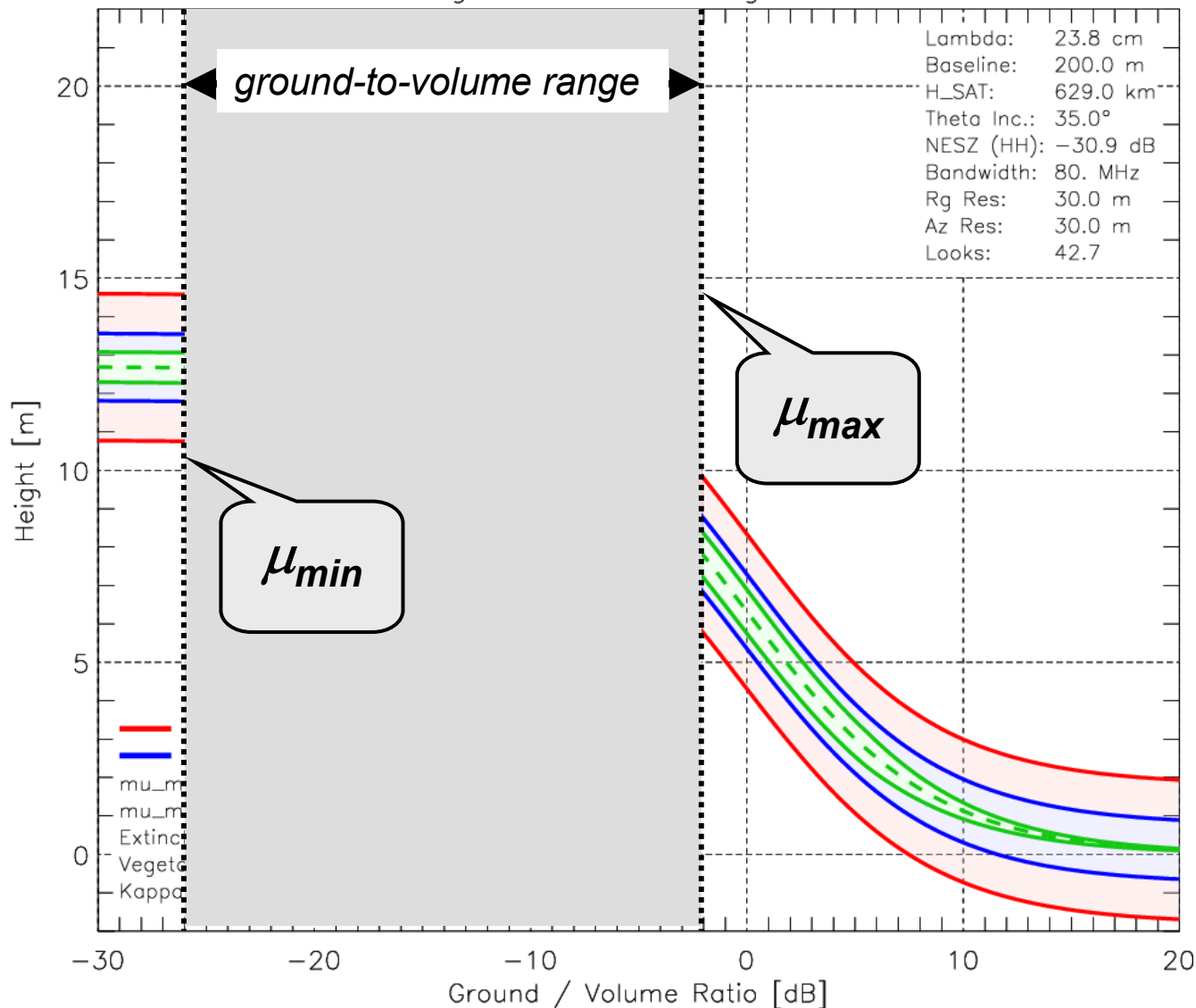
coherence ...

Height Error and Height Bias



- variation of phase center with different polarisations
- height errors due to volume decorrelation
- height errors without temporal decorrelation
- height errors with temporal decorrelation (here: $\gamma_{tmp}=0.9$)

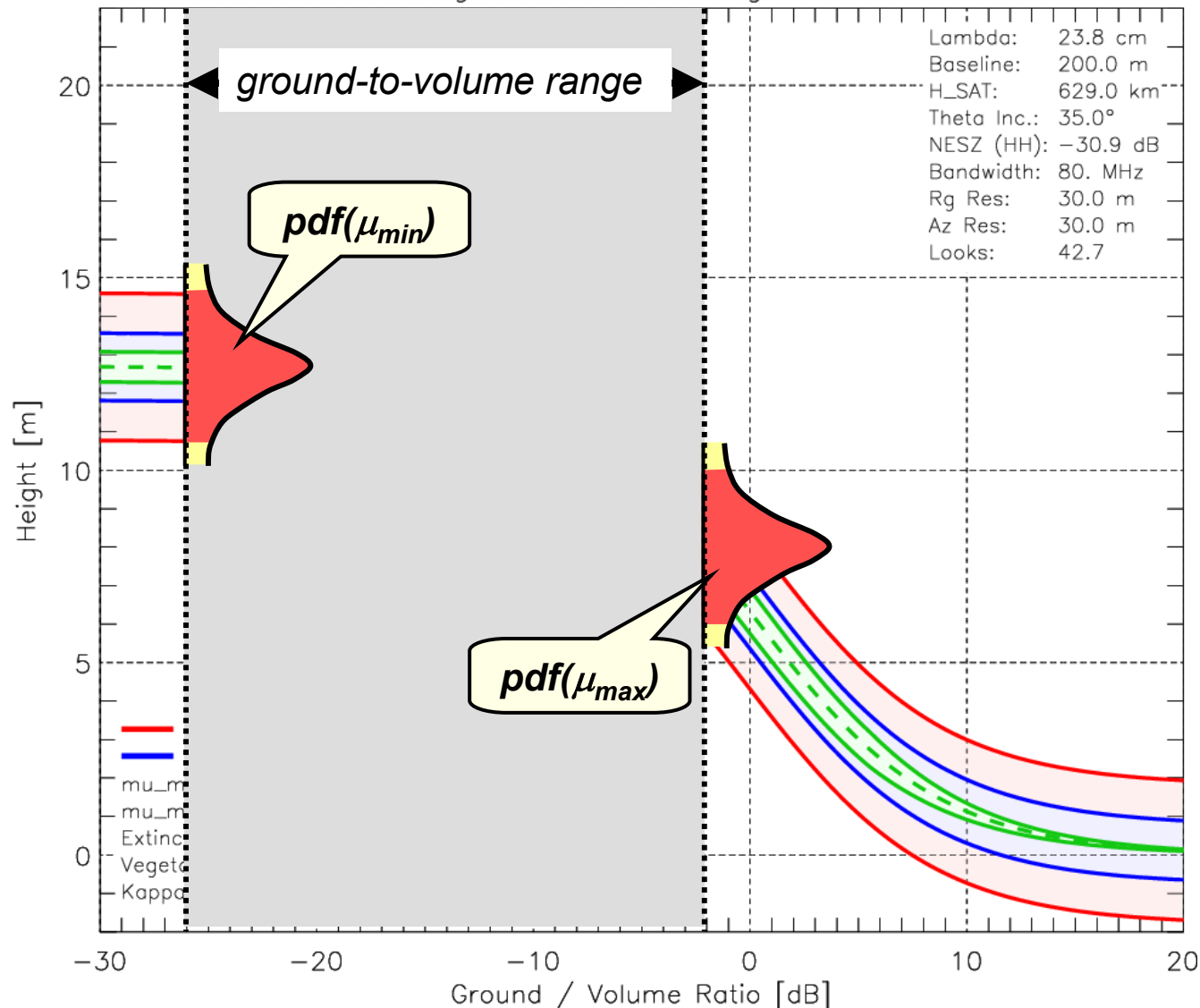
Height Error and Height Bias



*Accessible Range of
Ground-to-Volume
Scattering Ratios by
Different Polarisations:*

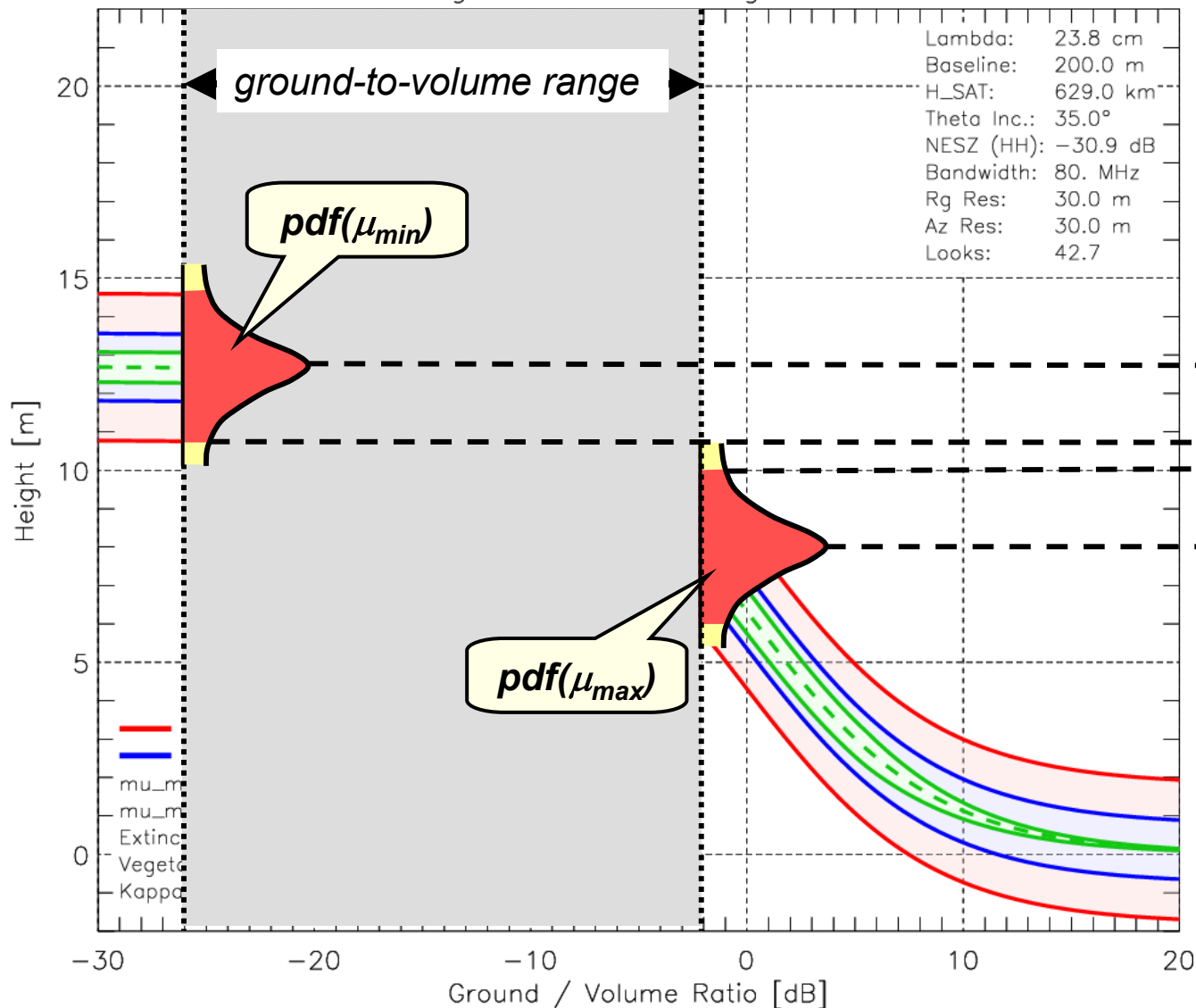
$$[\mu_{min}, \mu_{max}]$$

Height Error and Height Bias

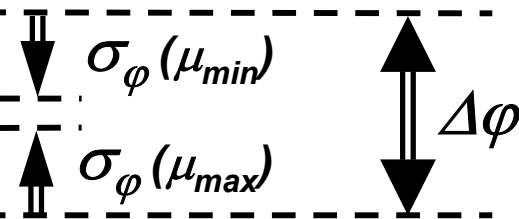


$pdf(\mu_{min})$ and $pdf(\mu_{max})$ describe the probability distribution of the phase center estimates at the two extremes of the ground-to-volume range μ_{min} and μ_{max} .

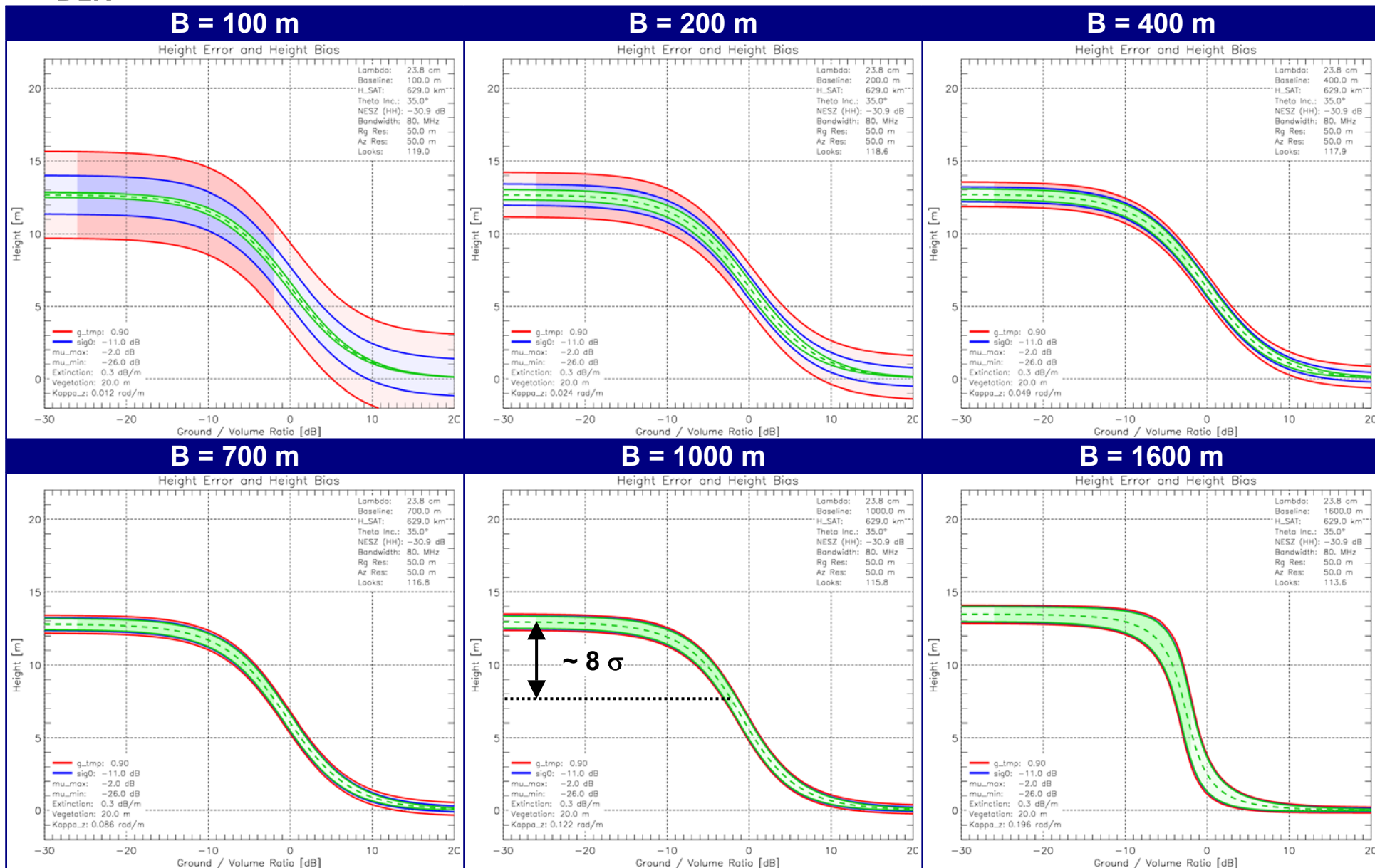
Height Error and Height Bias

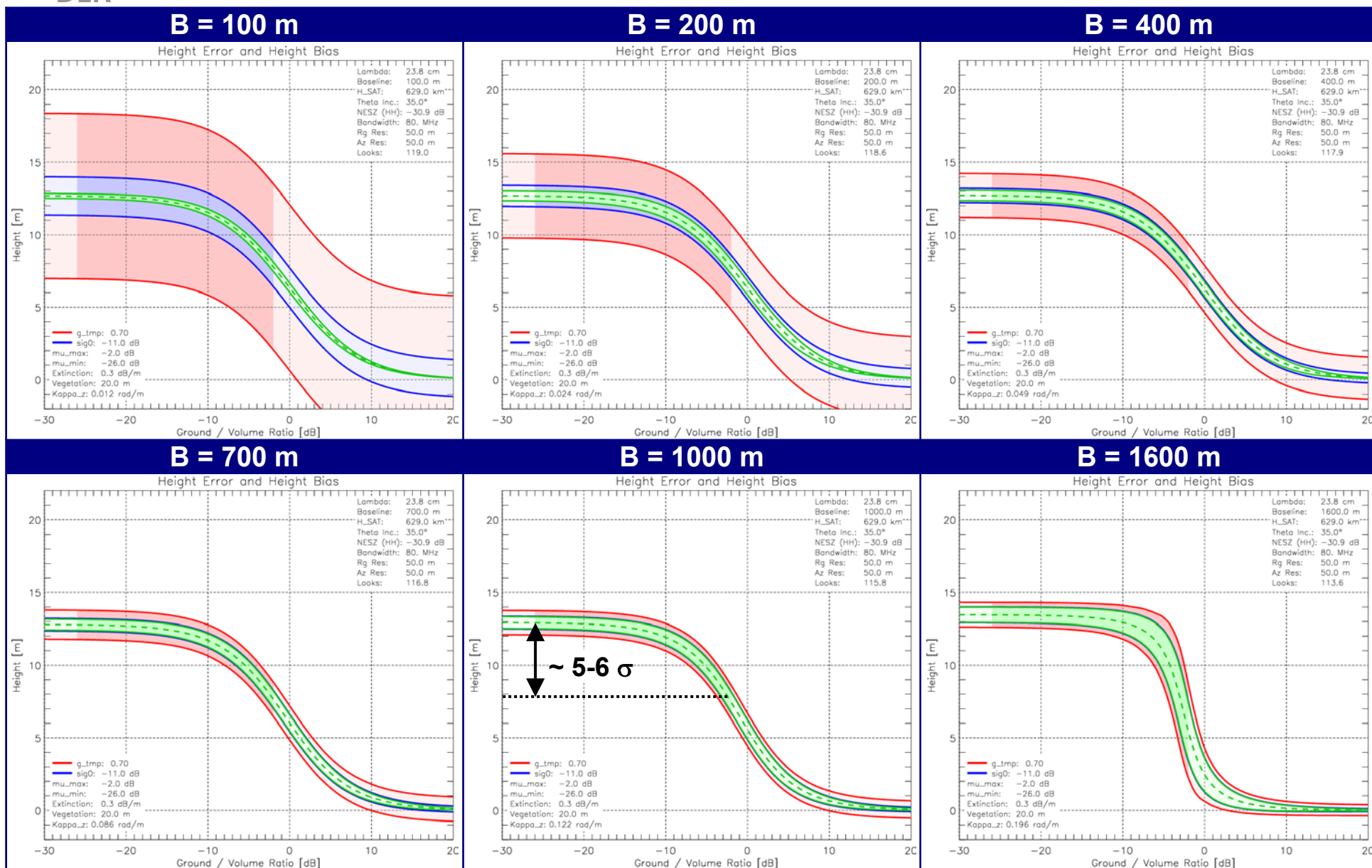


Separation of
Phase Centers:

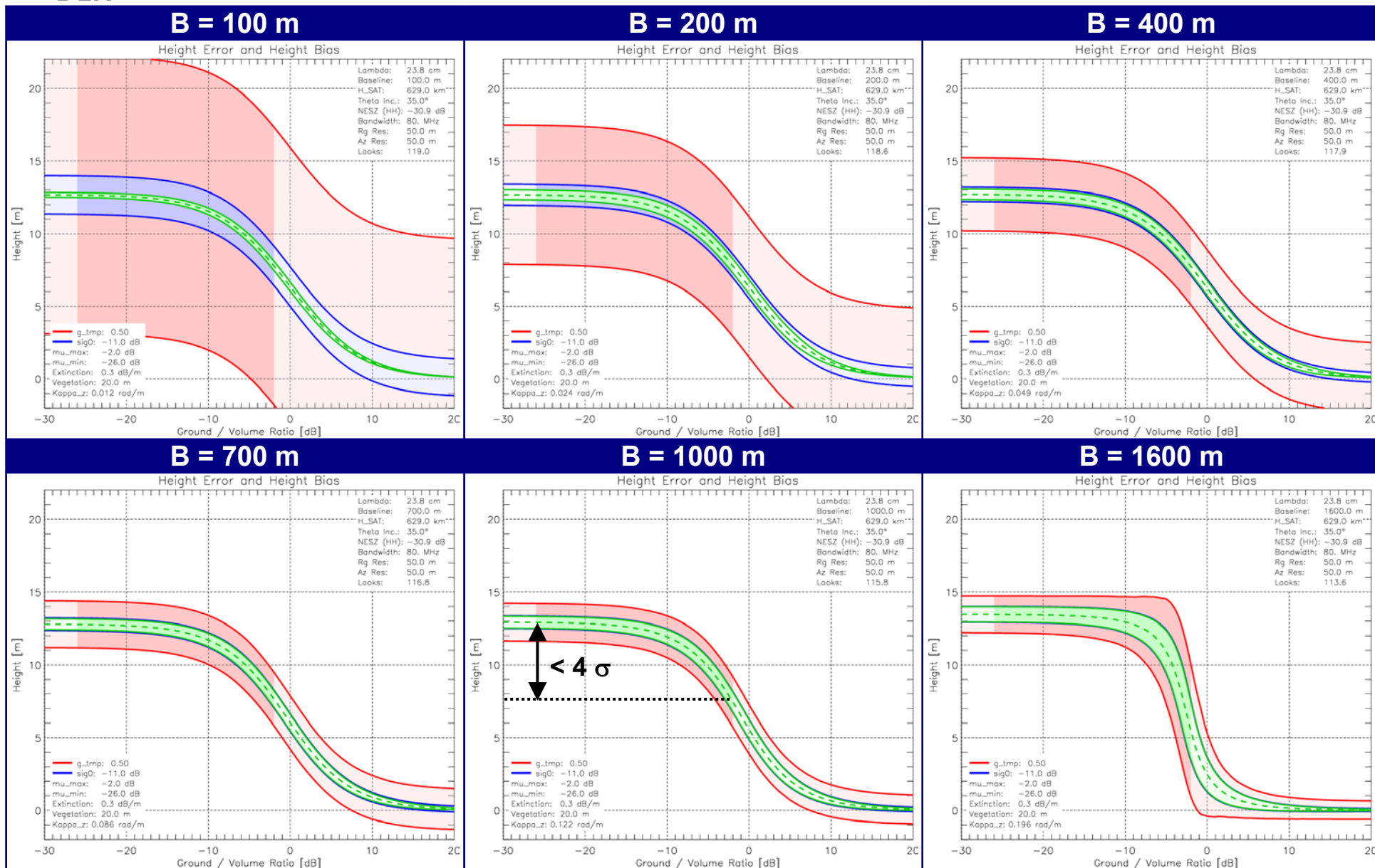


$$\Delta\phi \gg \sigma_{\phi}(\mu_{min}) + \sigma_{\phi}(\mu_{max})$$

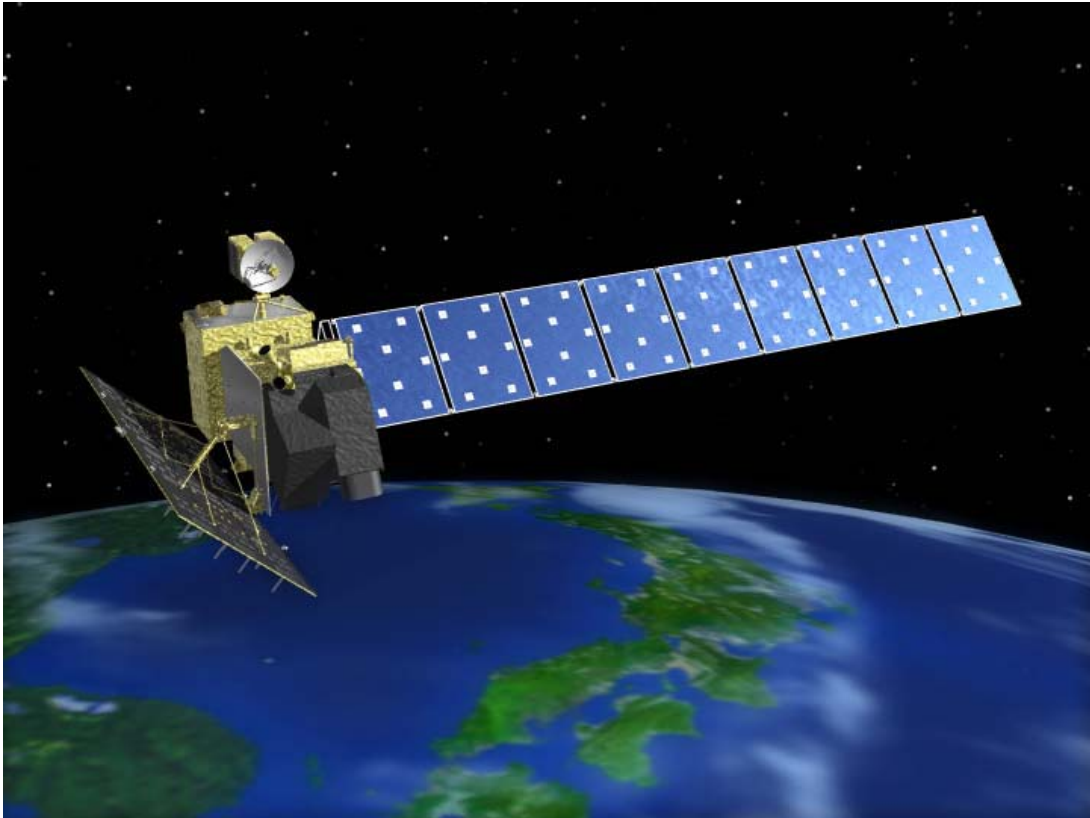


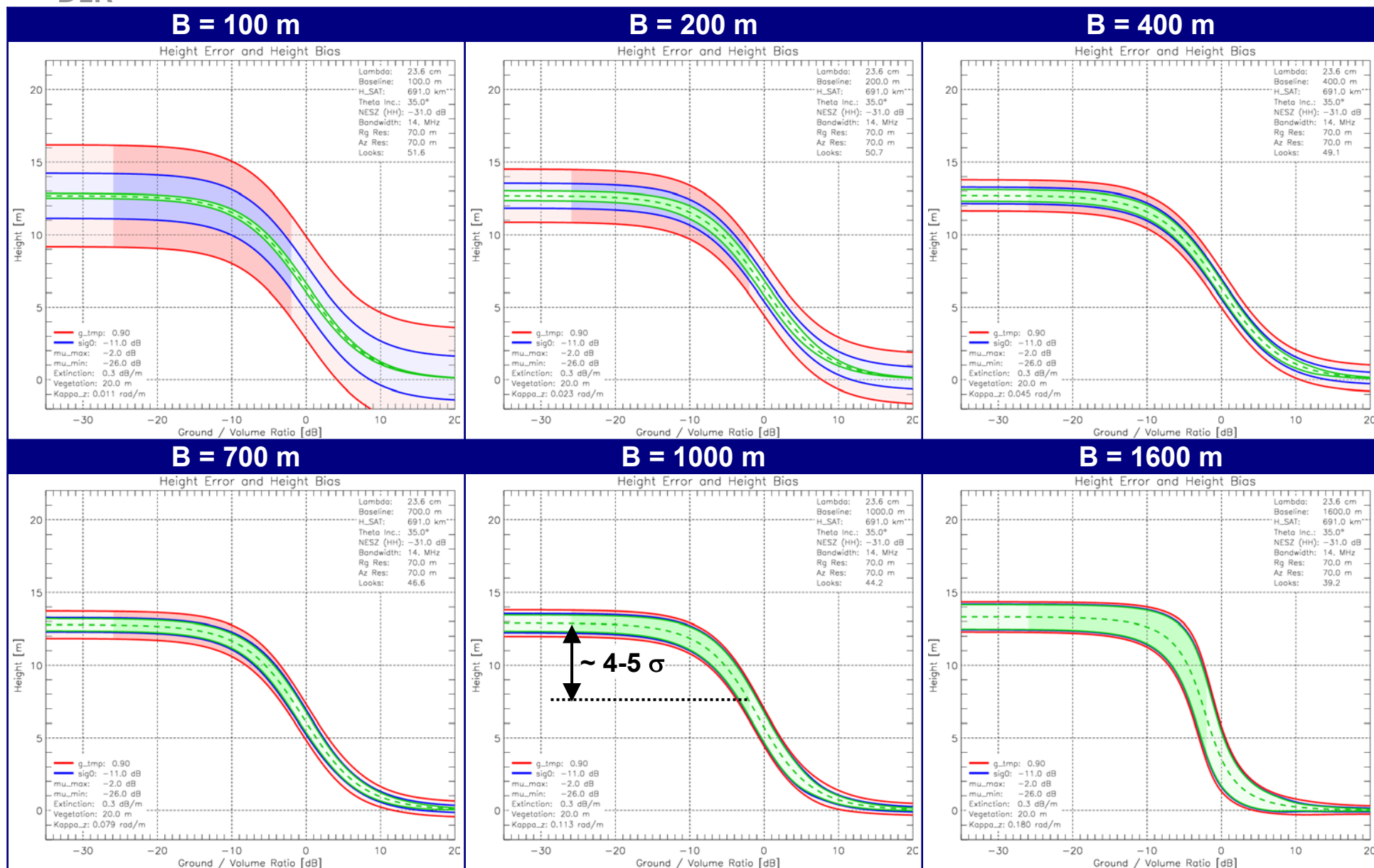


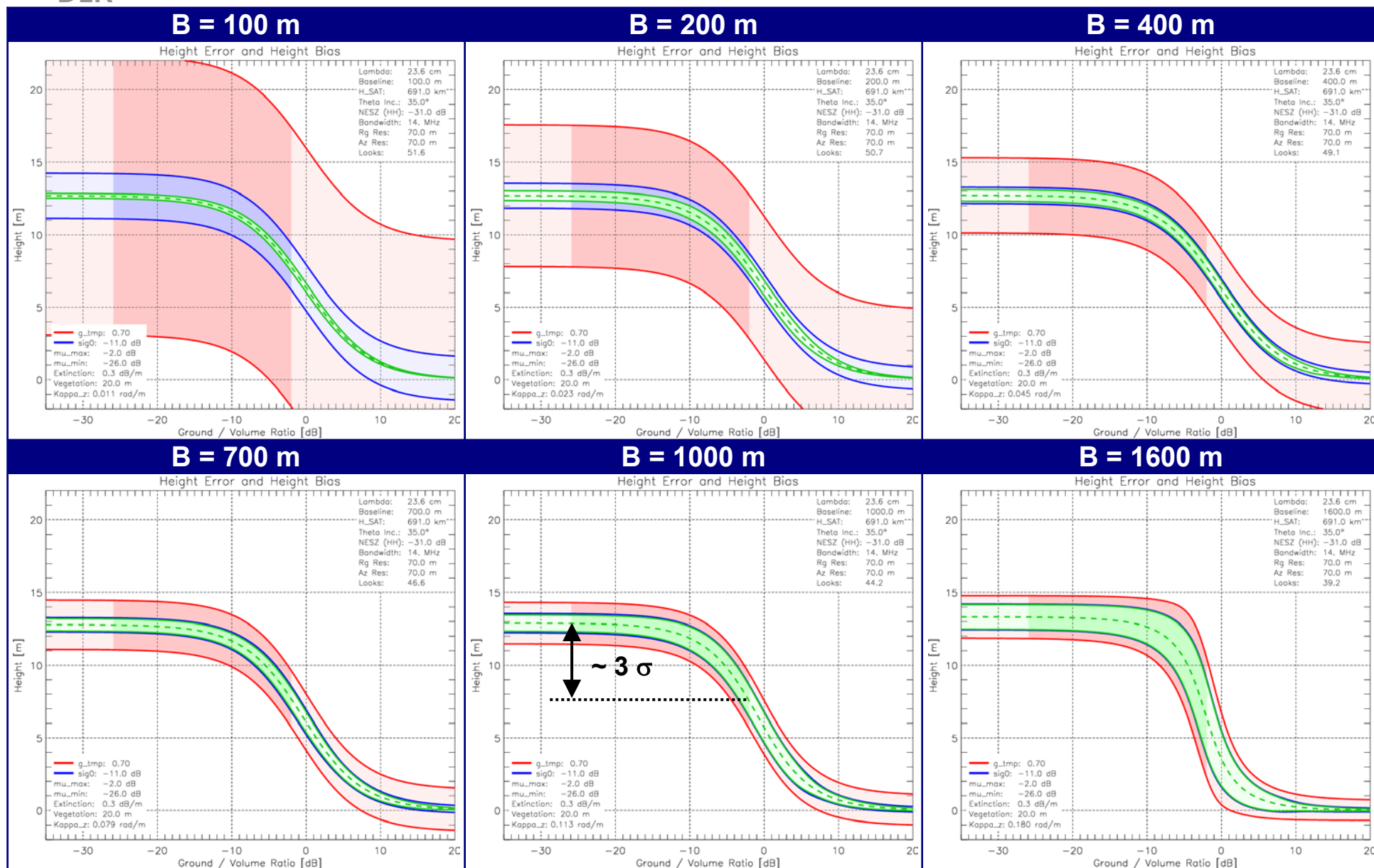
TerraSAR-L ($\gamma_{tmp} = 0.5$, 50 m x 50 m)

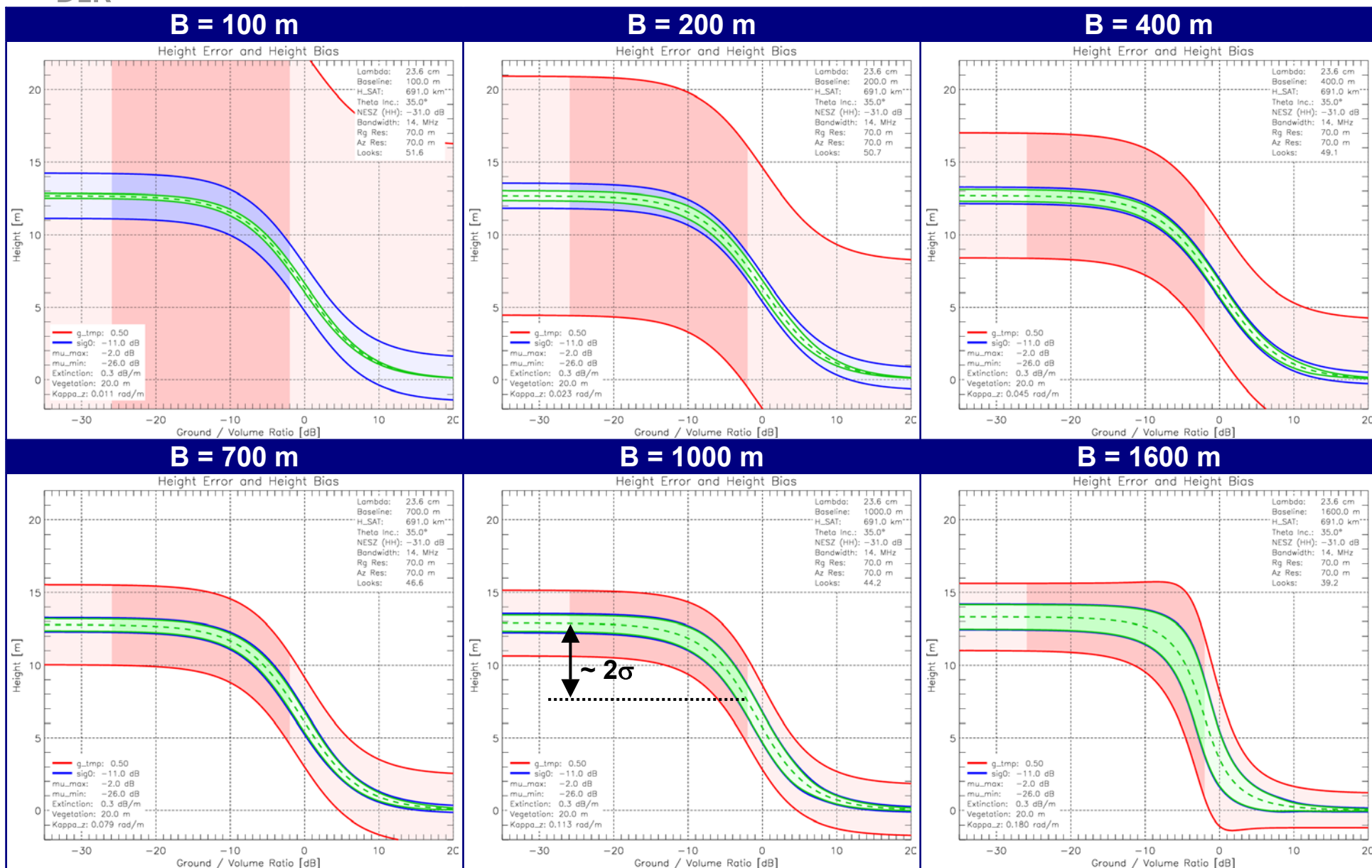


- A reliable separation of the phase centres will become possible for the Scots pine forest reference scenario at low to moderate temporal decorrelation ($\gamma_{temp} > 0.7$).
- Higher temporal decorrelation ($0.7 < \gamma_{temp} < 0.5$) may require an increase of the independent post spacing.
- Optimum performance is predicted for perpendicular baselines in the order of 700 m to 1000 m assuming a forest height of 20 m ($\kappa_z \sim 0.1$ rad/m).
- The performance will strongly depend on the accessible range of ground-to-volume ratios provided by the different polarisations. This range is expected to improve for lower incident angles.

Parameter	Value	
Wavelength	0.236 m	
Orbit Height	691 km	
Chirp Bandwidth	14 MHz (quad pol.)	
Peak Transmit Power	2 kW	
Duty Cycle	3,5 % (7 % / 2)	
Noise Figure	4 dB	
Rx. + Proc Losses	1 dB	
Losses across swath	< 3 dB	
Atmospheric Losses	1 dB	
Antenna Size (Tx , Rx)	8.9 m x 3.1 m	
Co-Registration Accuracy	1/10 pixel	
Quantisation	4 bit (BAQ)	

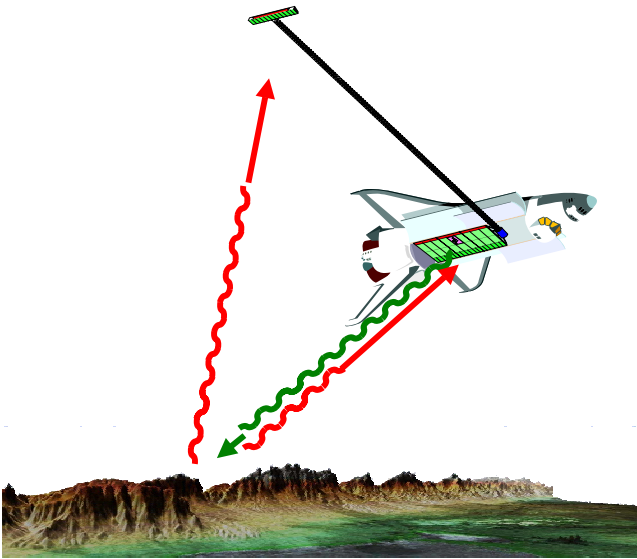
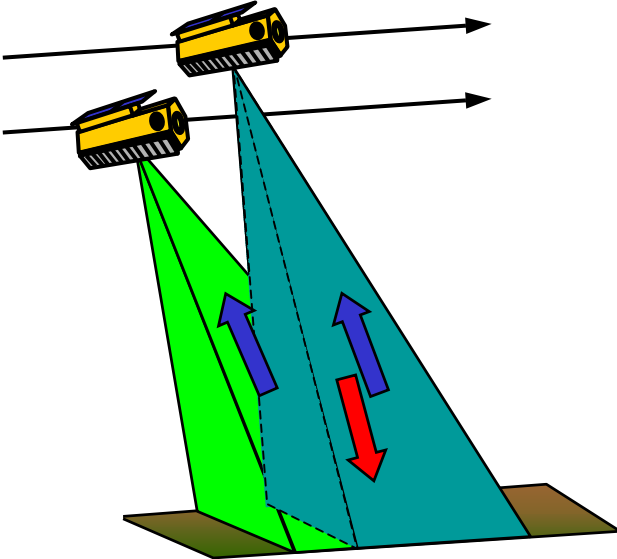
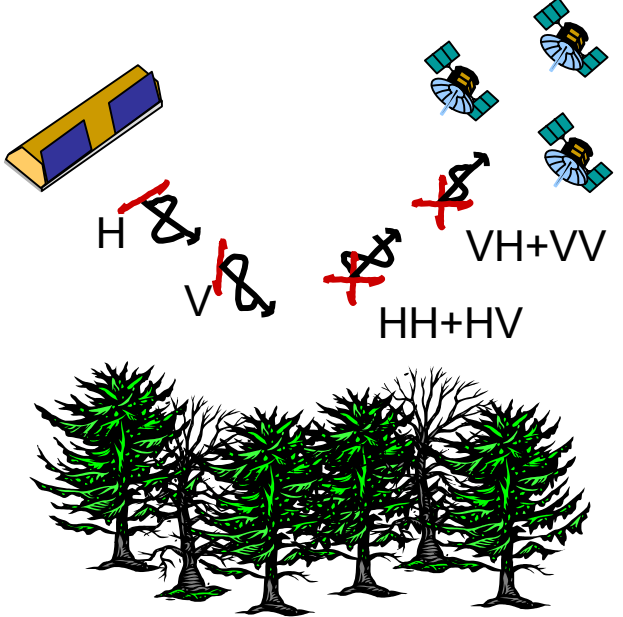




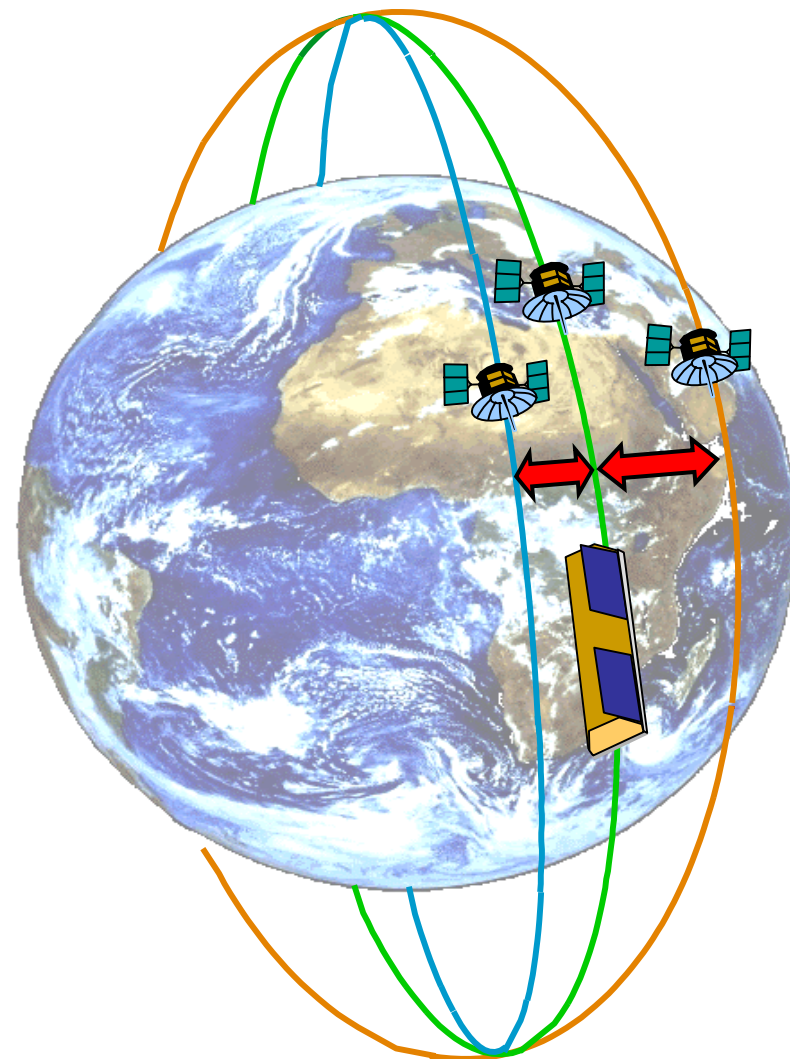


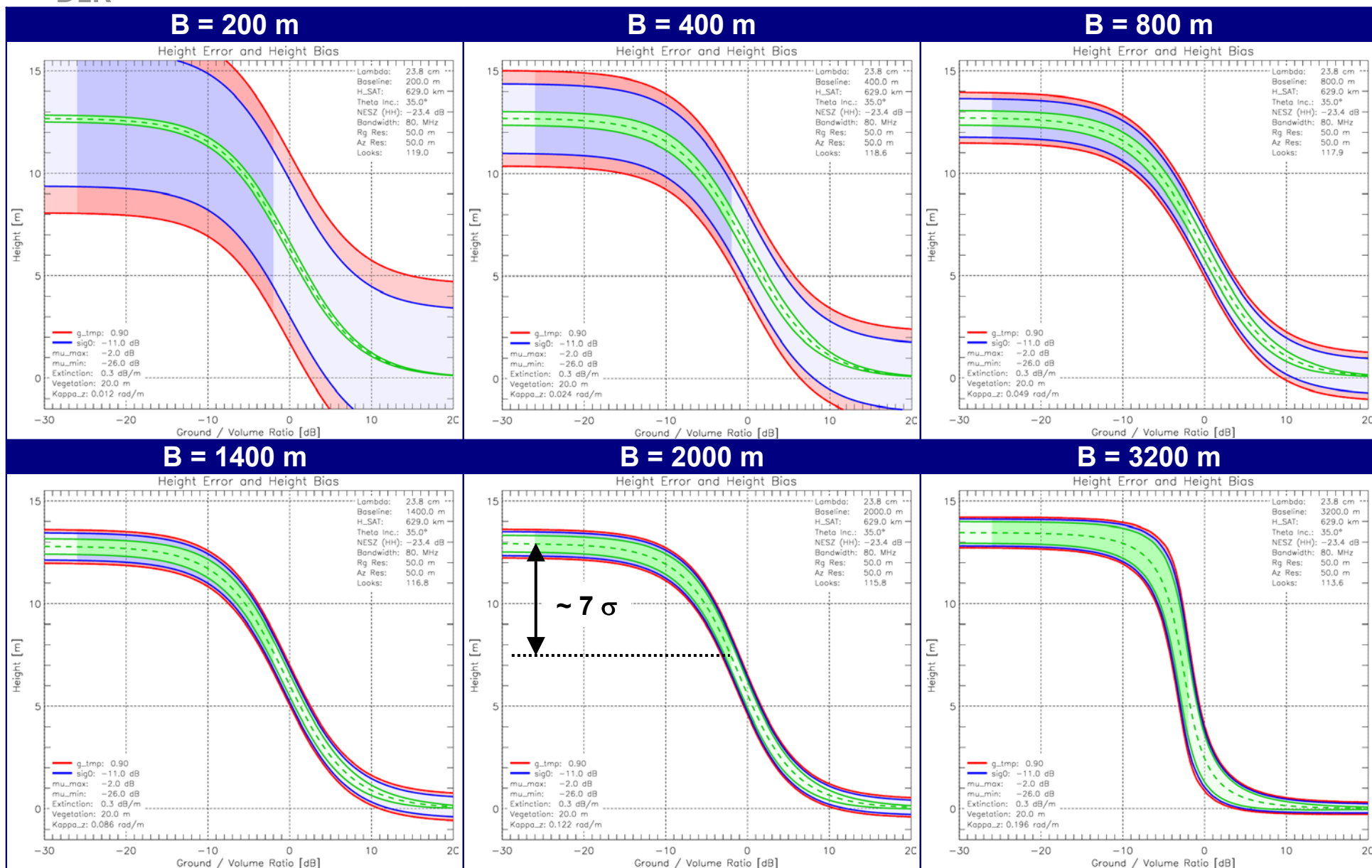
- The performance model predicts a rather poor PolInSAR inversion accuracy for the Scots pine reference scenario assuming an independent post spacing of 70 m x 70 m.
- The poor performance prediction is mainly due to the low number of independent looks resulting from the small system bandwidth of PALSAR (only 14 MHz for fully polarimetric operation).
- An acceptable performance could be achieved by increasing the independent post-spacing.
- ALOS has a rather long repeat cycle of 46 days. Hence, lower temporal coherence has to be expected as e.g. for TerraSAR-L with a repeat cycle of 14 days.

Single Pass Mission Scenarios

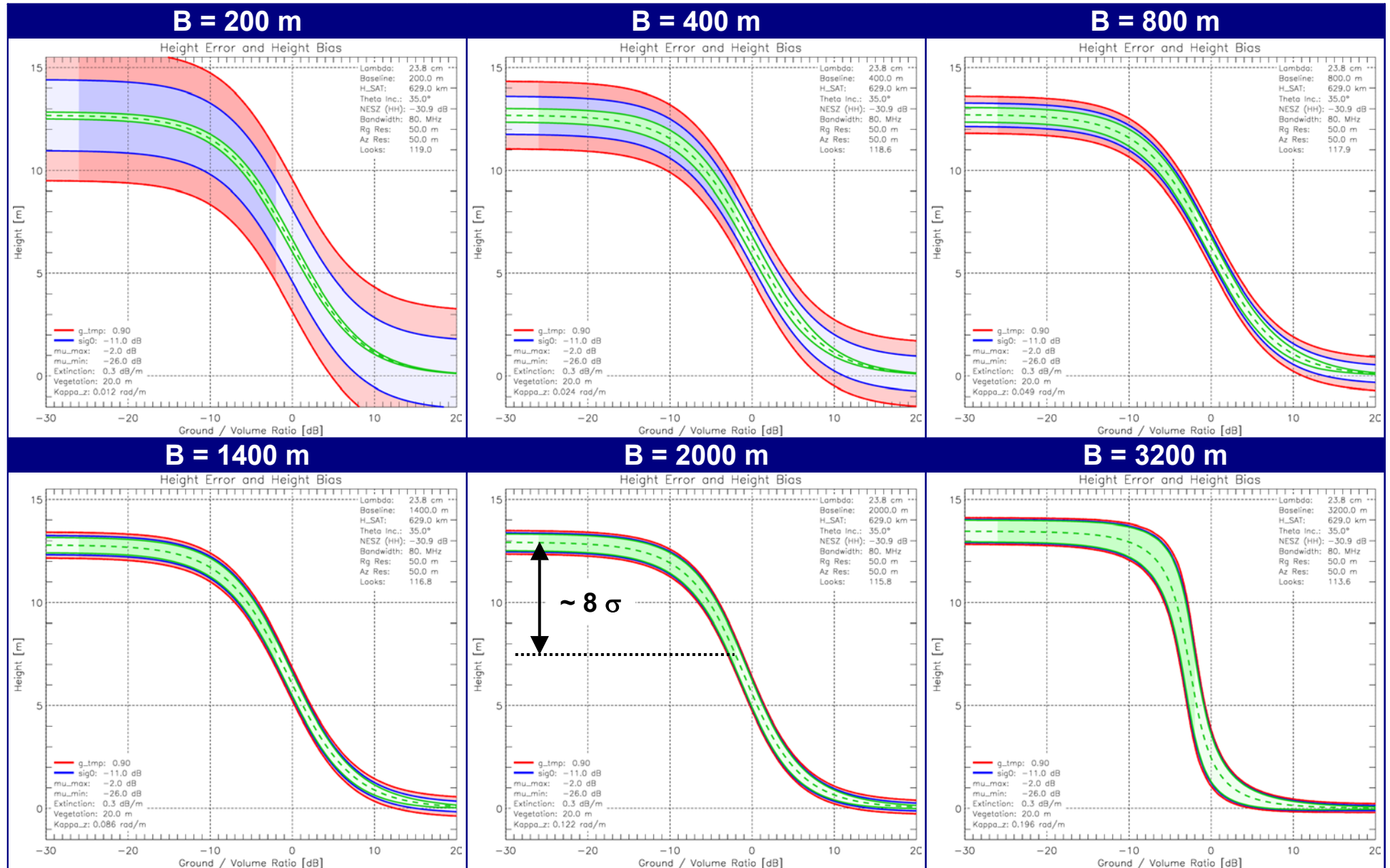
Satellite Boom Concepts (Precursor: SRTM)	Twin-Satellite Formations (TanDEM-X, Radarsat 2/3)	Multistatic SAR (Cartwheel, Pendulum, Voice)
		
<ul style="list-style-type: none"> •Maximum mast length (< 100 m) and minimum altitude (> 400 km) limit interferometric baseline •Ping-pong mode could increase interferometric baseline on the cost of higher ambiguities 	<ul style="list-style-type: none"> •Reconfigurable system for a wide range of interferometric baselines •High sensitivity •Flexible instrument operation •Phase synchronisation in ping/pong mode 	<ul style="list-style-type: none"> •Low cost passive receivers •Provides multiple interferometric baselines in a single pass •Increased susceptibility to ambiguities •Calibration more challenging

Parameter	Value
Wavelength	0.238 m
Orbit Height	629 km
Chirp Bandwidth	80 MHz
Peak Power (radiated)	4.4 kW
Duty Cycle (altern. pol.)	3,5 % (7 % / 2)
Noise Figure	2.5 dB (incl. swath)
Losses (proc., swath, ...)	< 6 dB
Antenna Size (Tx)	11 m x 2.86 m
Antenna Size (Rx)	circular: 3 m
Proc. Bandwidth	1200 Hz
Co-Registration Acc.	1/10 pixel
Quantisation	4 bit (BAQ)
Incident Angle	35°
Along-Track Displ. (Rx)	1 km
Antenna Size (Tx)	11 m x 2.86 m
Sigma Nought	> -11 dBm ² /m ² (HH,VV) > -14 dBm ² /m ² (HV,VH)
Post Spacing	50 m x 50 m
Extinction	0.3 dB/m
Volume Height	20.0 m





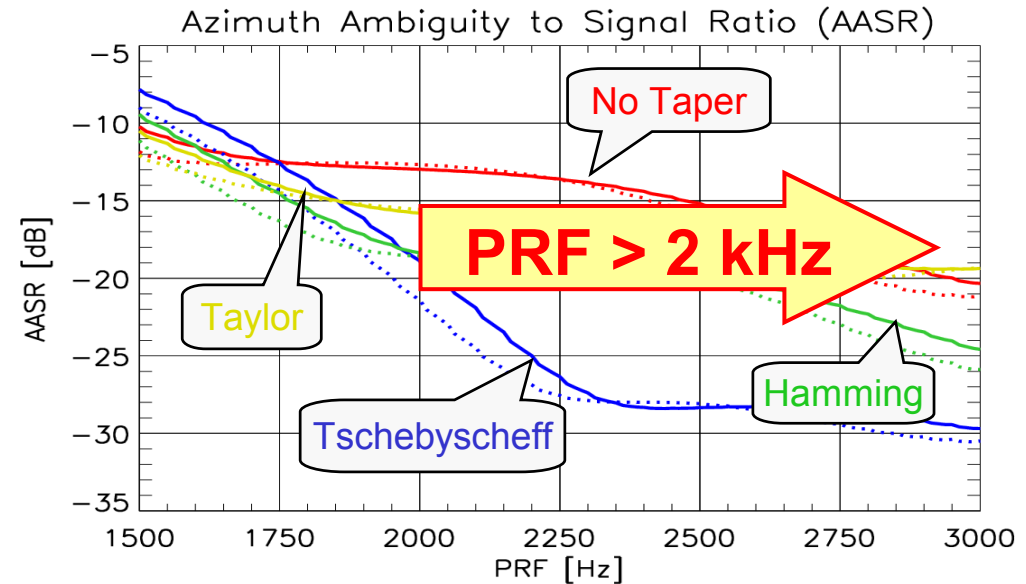
TerraSAR-L Tandem ($\gamma_{tmp} = 0.9$, 50 m x 50 m)



Critical Issues

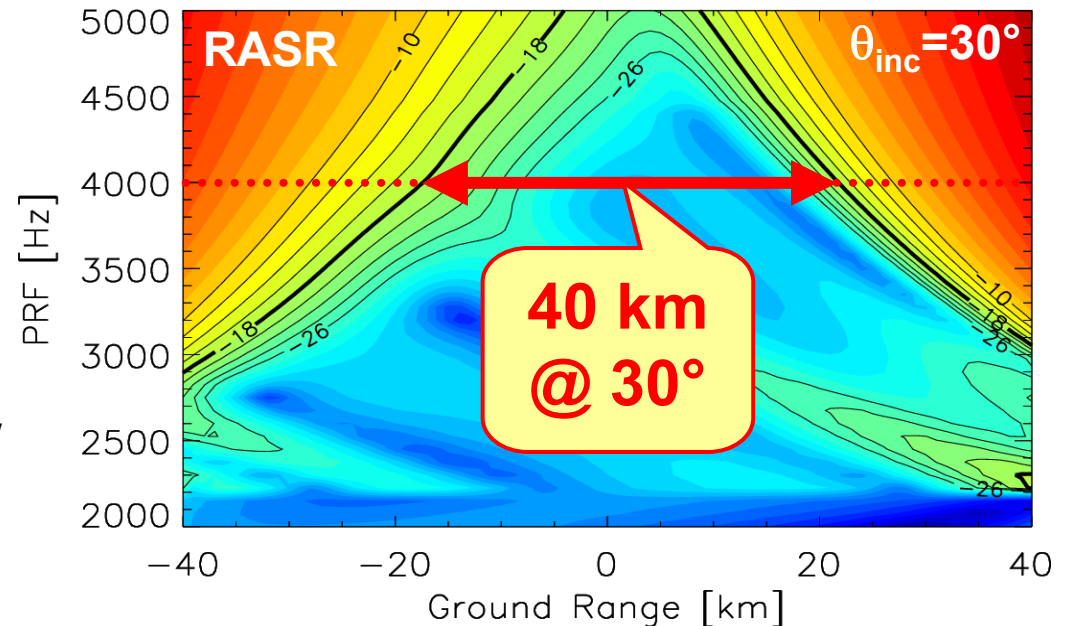
• Azimuth Ambiguities

- cartwheel-like configurations have an increased susceptibility to ambiguities due to the small Rx-antennas
- azimuth ambiguities in cartwheel like configurations may be reduced by an appropriate tapering of the transmit antenna



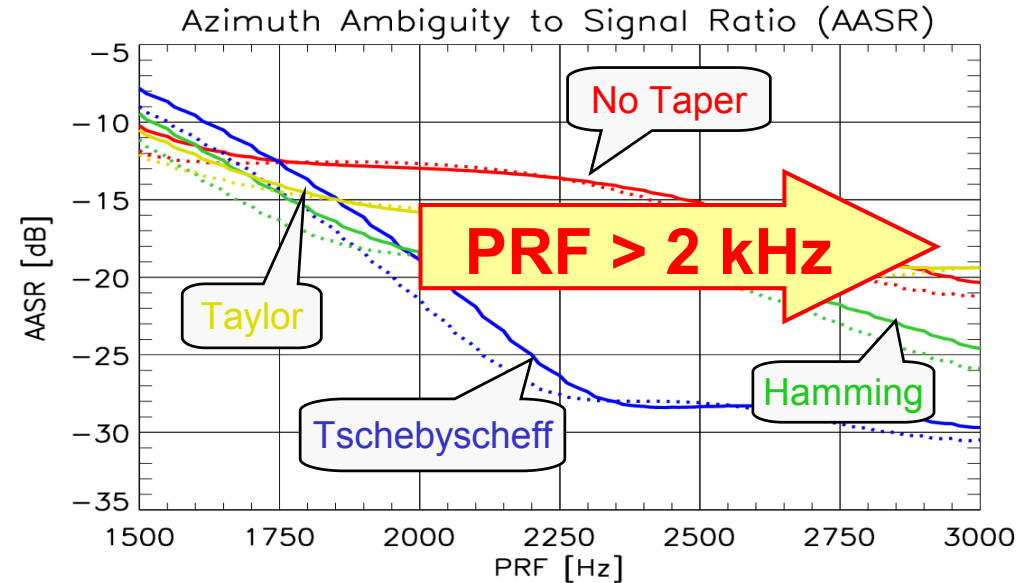
• Range Ambiguities

- a fully polarimetric mode requires alternating transmit polarisations which will increase the PRF by a factor of two
- range ambiguities will limit the unambiguous swath width at shallow incident angles (e.g. 40 km @ 30° and 12 km @ 35° in TSL-CW)



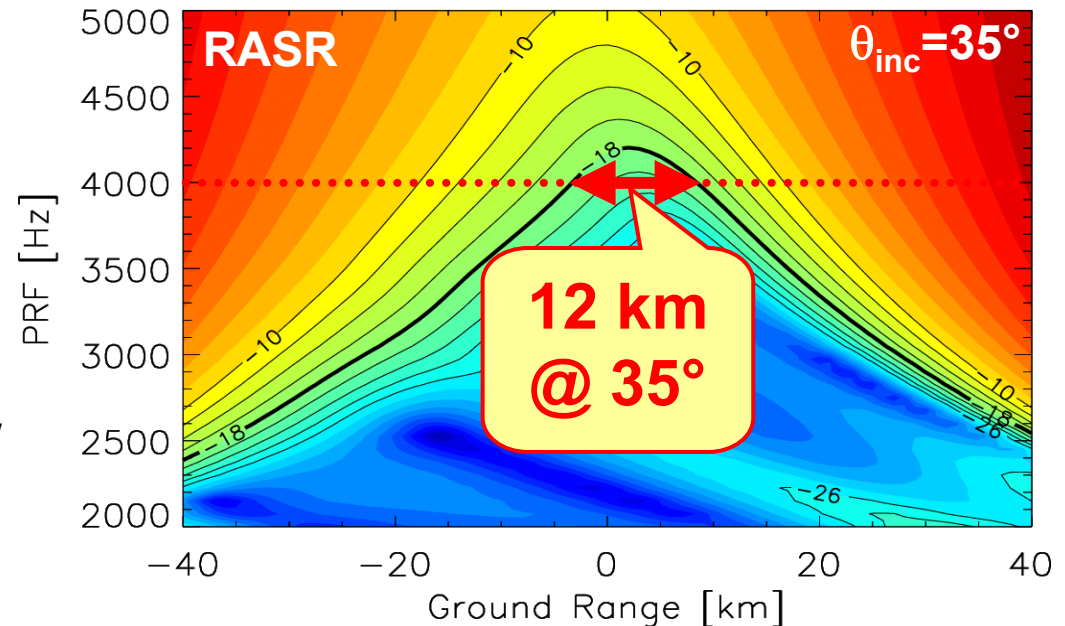
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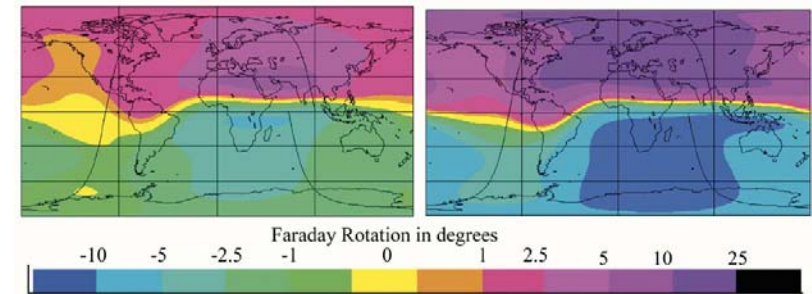
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• **Faraday Rotation**

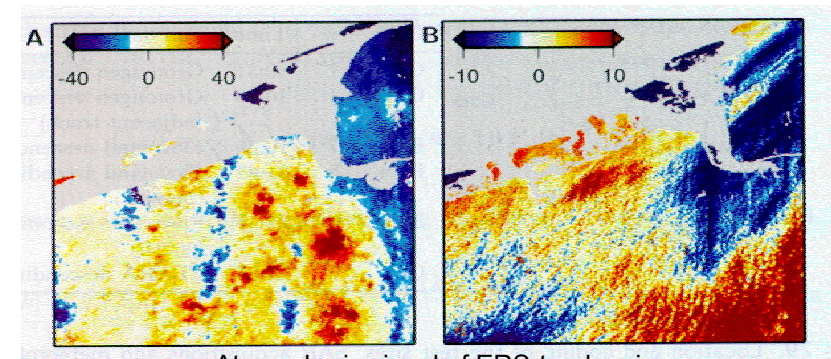
- may become several 10th of degrees (L-band, solar max.)
 - *single pass*: equal rotation for all channels (less critical for most PolInSAR applications)
 - *repeat pass*: differential rotation within interferometric channels
- correction of Faraday rotation possible within 3° to 5° (cf. Freeman, 2004)
- open issues: (1) tolerance level for different PolInSAR applications
(2) bistatic angle in cartwheel configurations



Predictions of one-way Faraday rotation in L-Band for low (left) and high (right) solar activity (P. Wright, 2003)

• **Atmospheric Disturbances**

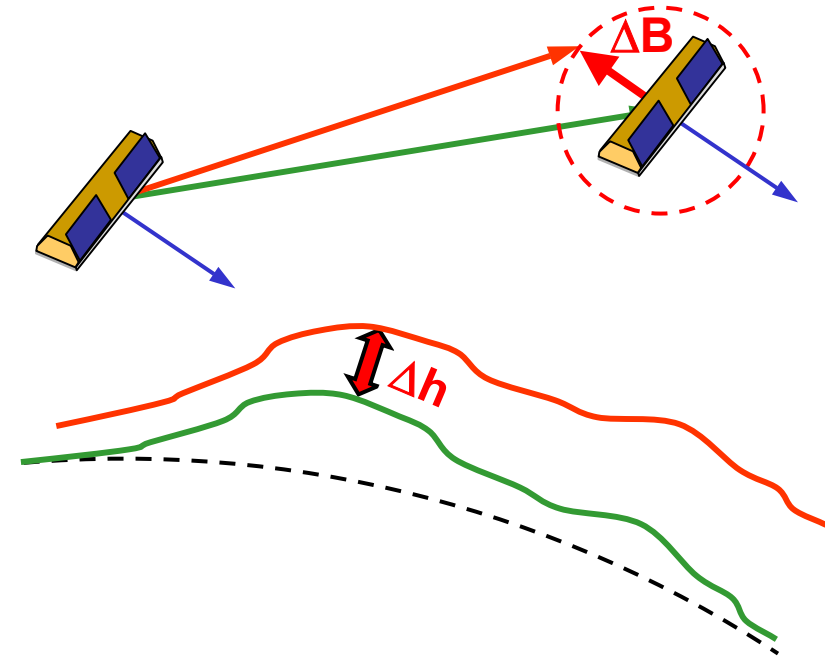
- different atmospheric conditions may cause phase disturbances in repeat pass InSAR
- low frequency bias of ground topography phase, residual PolInSAR parameters are less affected
- becomes more severe for shorter wavelengths



Atmospheric signal of ERS-tandem in millimeters zenith delay from (R. Hanssen, 2001)

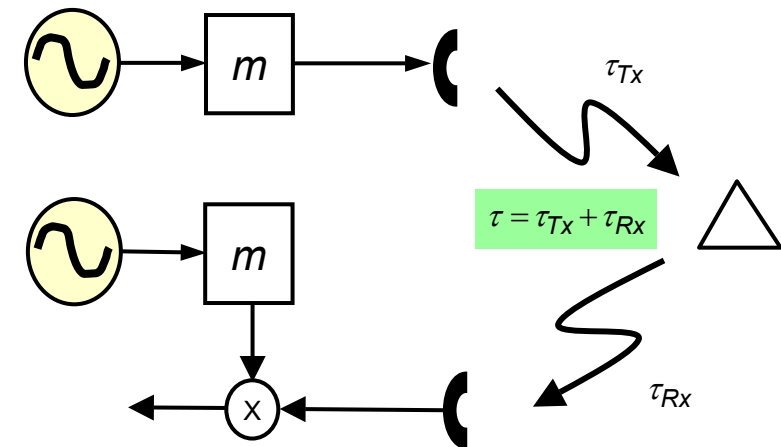
• **Baseline Estimation Errors**

- will mainly affect estimates of the absolute ground topography phase φ_0
- may also introduce a slight bias in the inversion process due to the use of a wrong $\Delta\theta$
- single-pass configurations enable precise estimation of baseline vector with mm accuracy based on relative GPS carrier phase evaluations



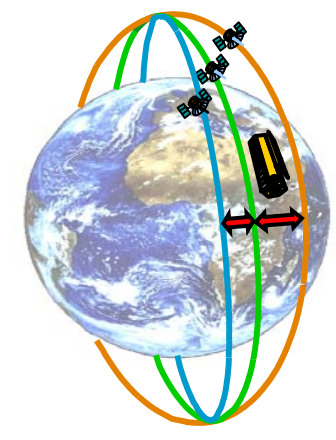
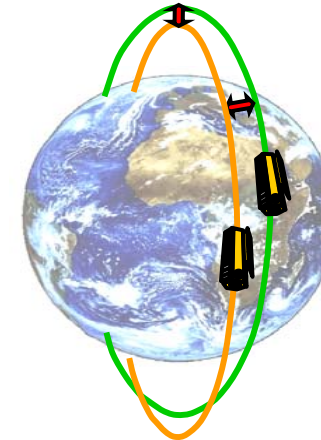
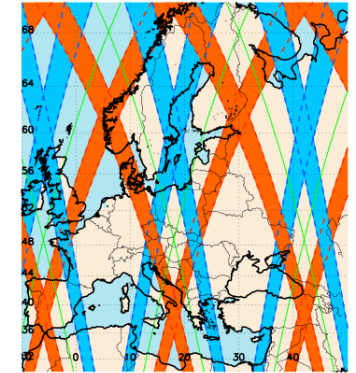
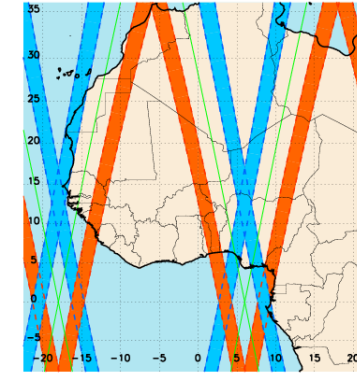
• **Oscillator Phase Errors**

- mainly affect single pass configurations in bistatic mode
- uncorrelated oscillator noise will introduce low frequency phase errors in azimuth
- will mainly affect estimates of the absolute ground topography phase φ_0



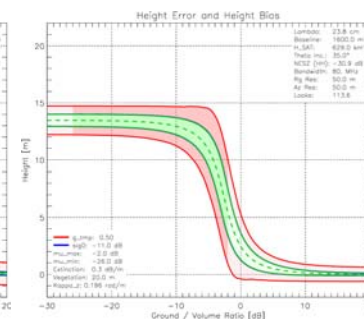
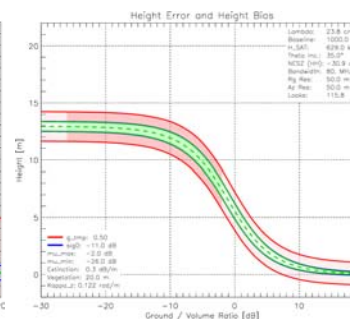
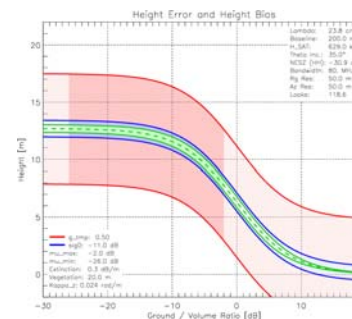
• Orbit Selection

- repeat pass missions require short repeat cycles to minimize temporal decorrelation
- short repeat cycles cause large gaps between the ground tracks, thereby requiring a wide range of incident angles to achieve global coverage
- single pass missions require appropriately designed satellite formations to provide the desired baselines and to avoid any collision risk



• Baseline (κ_z) Optimisation

- the PolInSAR performance depends strongly on the chosen interferometric baseline
- system optimization requires some a priori knowledge about the imaged scene or an acquisition with multiple baselines



• Temporal Decorrelation and Multi-Looking

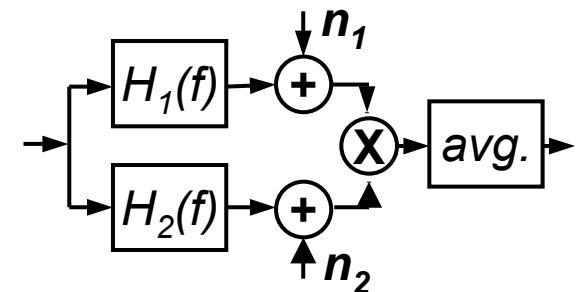
- the current performance analysis uses the multiplicative coherence model, i.e. temporal errors are modelled as additive white Gaussian noise in each SAR image
- is such a model appropriate ? (precipitation and wind may also cause a local shift of the phase centres)
- should we hence expect correlated errors between adjacent resolution cells ? Will multi-looking reduce temporal phase errors in the predicted manner ?
- Are all polarisations affected in the same way ? (ground and volume may be affected differently)

• Bistatic Calibration

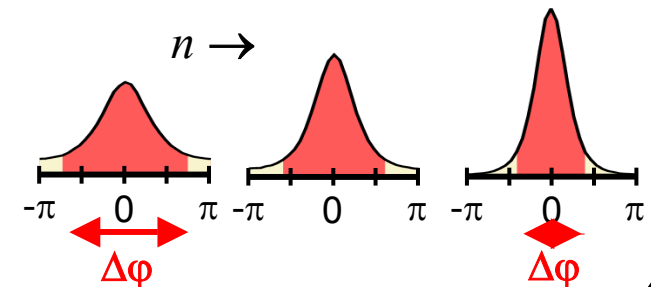
- bistatic acquisitions may require new calibration techniques (e.g. due to invisible corner reflectors for large bistatic angles)

$$\gamma_{tot} = \gamma_{temp} \cdot \gamma_{SNR} \cdot \gamma_{vol} \cdots$$

$$\gamma_{SNR} = \frac{1}{\sqrt{1 + SNR_1^{-1}} \cdot \sqrt{1 + SNR_2^{-1}}}$$



$$pdf_{\varphi}(\varphi) = f(\gamma, n)$$



- Repeat pass missions will suffer from significant deteriorations due to temporal decorrelation.
- The application area of boom interferometers is very limited due to the short baseline length.
- Satellite formations avoid temporal decorrelation and offer the opportunity to acquire interferometric data with arbitrary baseline lengths in a highly flexible and reconfigurable geometry.
- No show stopper has been found for the implementation of a multi-satellite single pass PolInSAR mission, but a careful mission design will be required.